

Erken B-türü β Cephei değişen yıldızı Gamma Pegasi'nın YTD (Yerel Termodynamik Denge) ve YTDO (Yerel Termodynamik Denge Olmayan) CNO bolluk sonuçları:

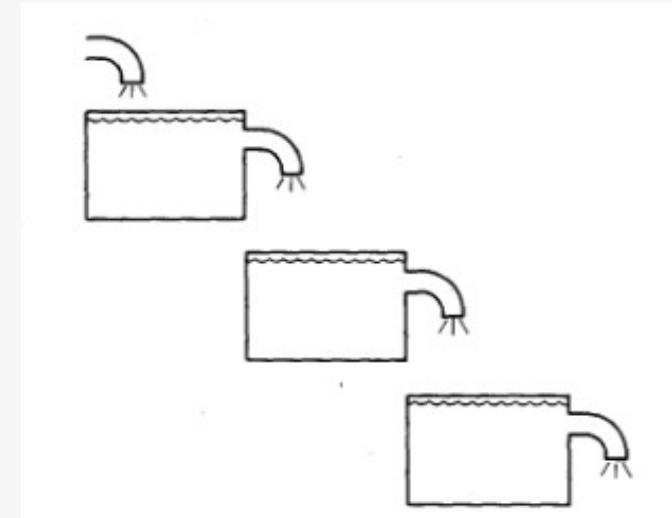
Taner Tanrıverdi¹

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**ULUSAL ASTRONOMİ KONGRESİ, 5-9 EYLÜL 2016, ATATÜRK
ÜNİVERSİTESİ, ERZURUM**

YTD (Yerel Termodinamik Denge)

- - Hidrostatik denge
 - Radyatif Denge
 - Termodinamik Denge
 - Yük Dengesi



¹ H 13.60	
³ Li 5.39 75.64	⁴ Be 9.32 18.21
¹¹ Na 5.14 47.29	¹² Mg 7.65 15.03
¹⁹ K 4.34 31.63	²⁰ Ca 6.11 11.87
³⁷ Rb 4.18 26.05	³⁸ Sr 5.70 11.03
⁵⁵ Cs 3.89 25.14	⁵⁶ Ba 5.21 10.00
⁸⁷ Fr 3.00 20.00	⁸⁸ Ra 5.28 10.15

Periodic System:
first and second ionization potential (eV)

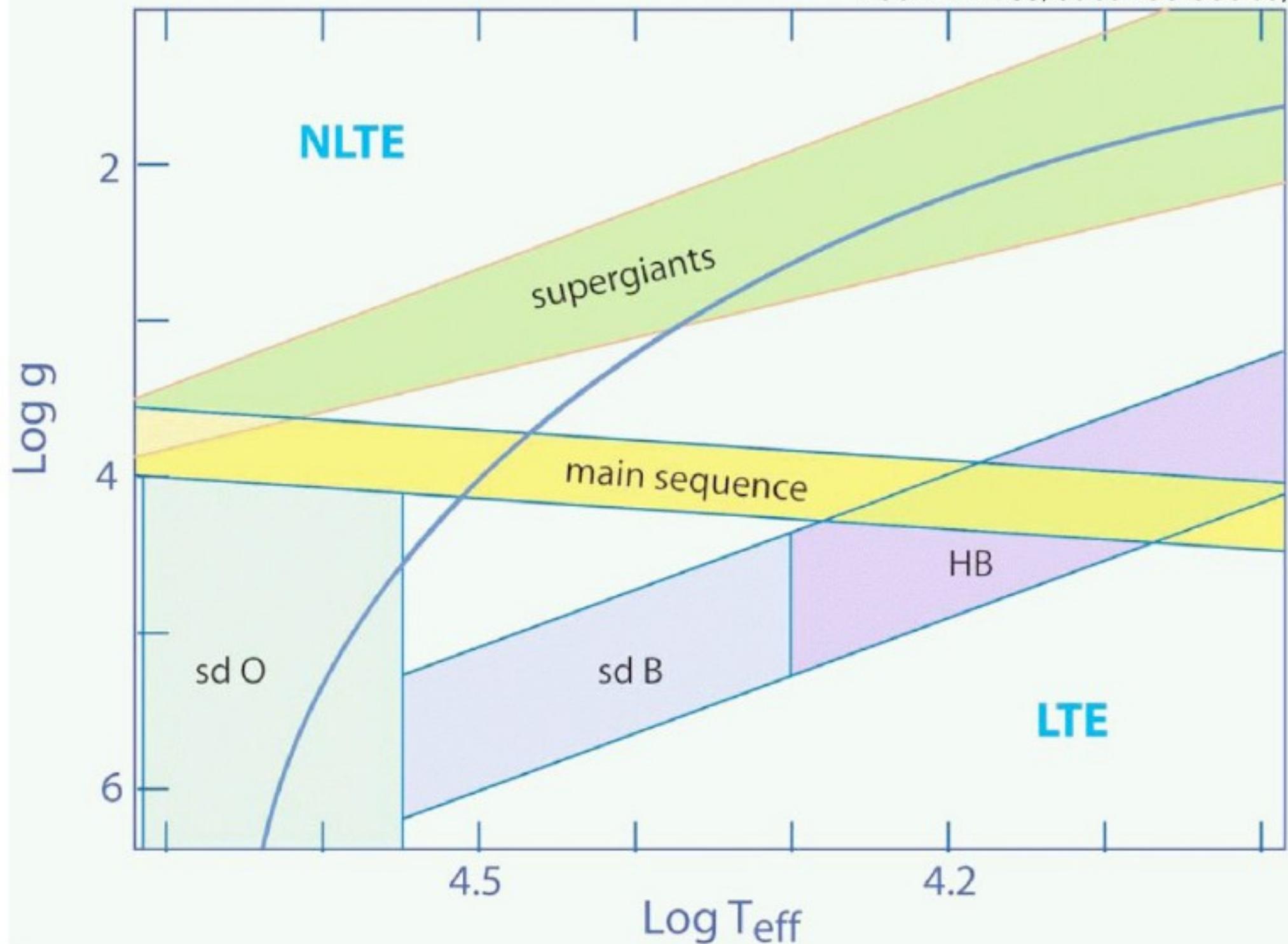
NLTE

LTE

					² He 24.59 54.42
⁵ B 8.30 25.16	⁶ C 11.26 24.38	⁷ N 14.53 29.60	⁸ O 13.62 35.12	⁹ F 17.42 34.97	¹⁰ Ne 21.57 40.96
¹³ Al 5.99 18.83	¹⁴ Si 8.15 16.35	¹⁵ P 10.49 19.73	¹⁶ S 10.36 23.33	¹⁷ Cl 12.97 23.81	¹⁸ Ar 15.76 27.63
³¹ Ga 6.00 20.51	³² Ge 7.90 15.94	³³ As 9.81 18.59	³⁴ Se 9.75 21.19	³⁵ Br 11.81 21.80	³⁶ Kr 14.00 24.36
⁴⁹ In 5.79 18.87	⁵⁰ Sn 7.34 14.63	⁵¹ Sb 8.64 16.53	⁵² Te 9.01 18.60	⁵³ I 10.45 19.13	⁵⁴ Xe 12.13 21.21
⁸¹ Tl 6.11	⁸² Pb 7.42	⁸³ Bi 7.29	⁸⁴ Po 8.42	⁸⁵ At 9.00	⁸⁶ Rn 10.75

Atmosfer Modelleri

- LTE
 - ATLAS9
 - MARCS
- NLTE
 - DETAIL
 - ATA
 - TLUSTY



- Erken Tür ve Süperdevler
 - NLTE
- $\geq A8$ tayf türü
 - Bańç Tersinirliği(Pressuure Inv.)
 -
- Soğuk Yıldızlar
 - F,G,K yıldızları

Model Atmosfer Programı

- **TLUSTY**
 - $27500 \text{ K} \leq \text{Teff} \leq 55000 \text{ K}$
 - $3.0 \leq \log g \leq 4.75$
 - Düzlem paralel

ELODIE

- 4100-6800 Å
- R~42000 ve 10000
- S/N~300

HD886

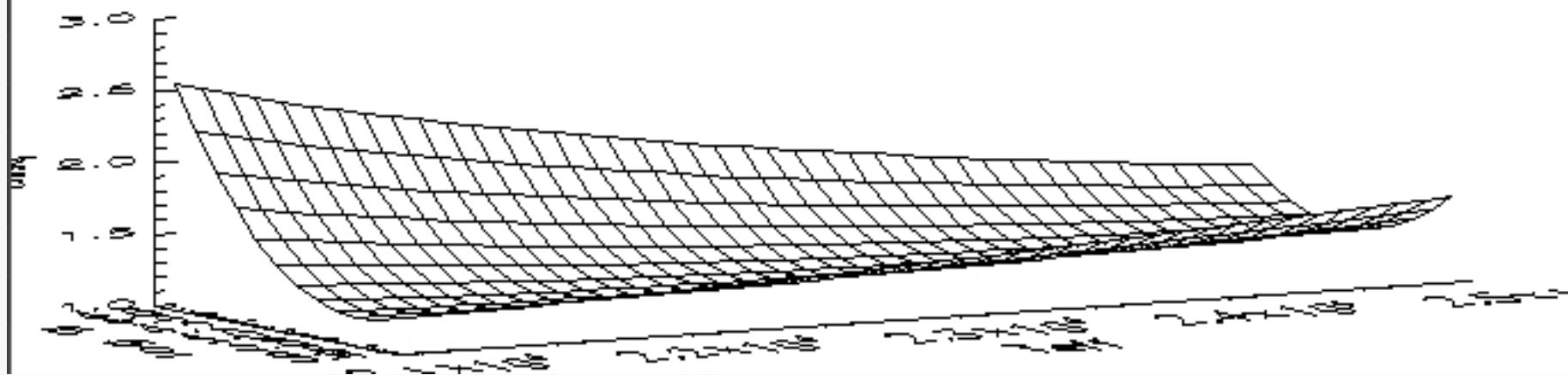
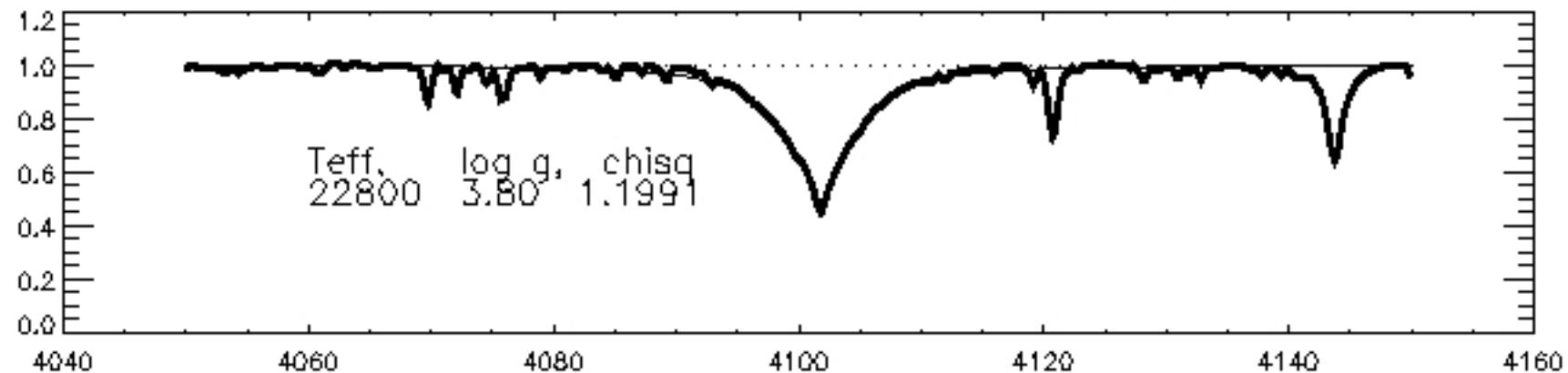
- B2 IV tayf türü
- Beta Cep değişen
 - Galaktik
 - $|l \sim 109$, $b \sim -46$
- Nieva&Przybilla,2014,A&A,566,7
- Gies&Lambert,1992,ApJ,387,673

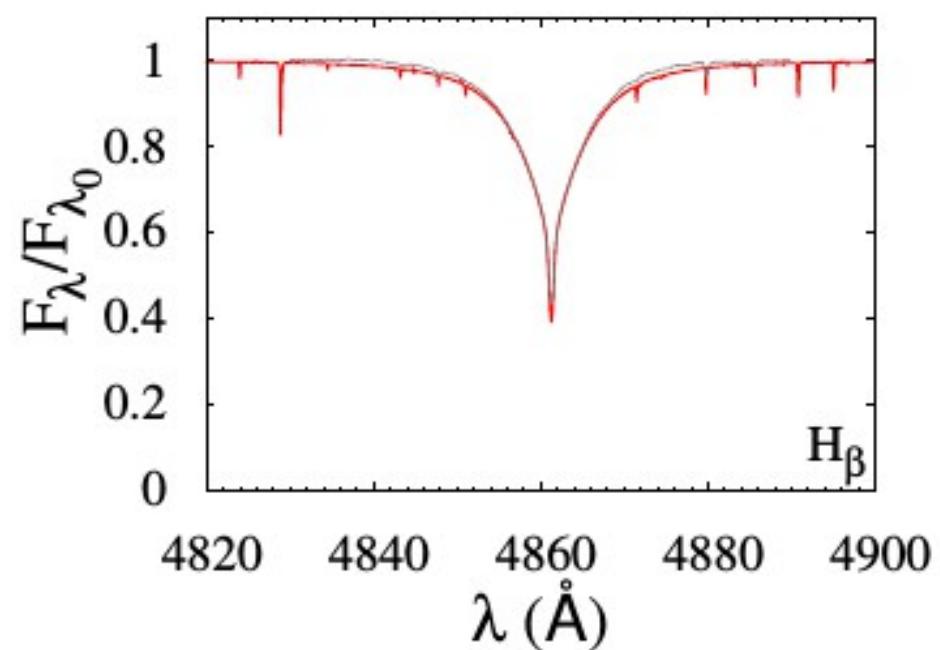
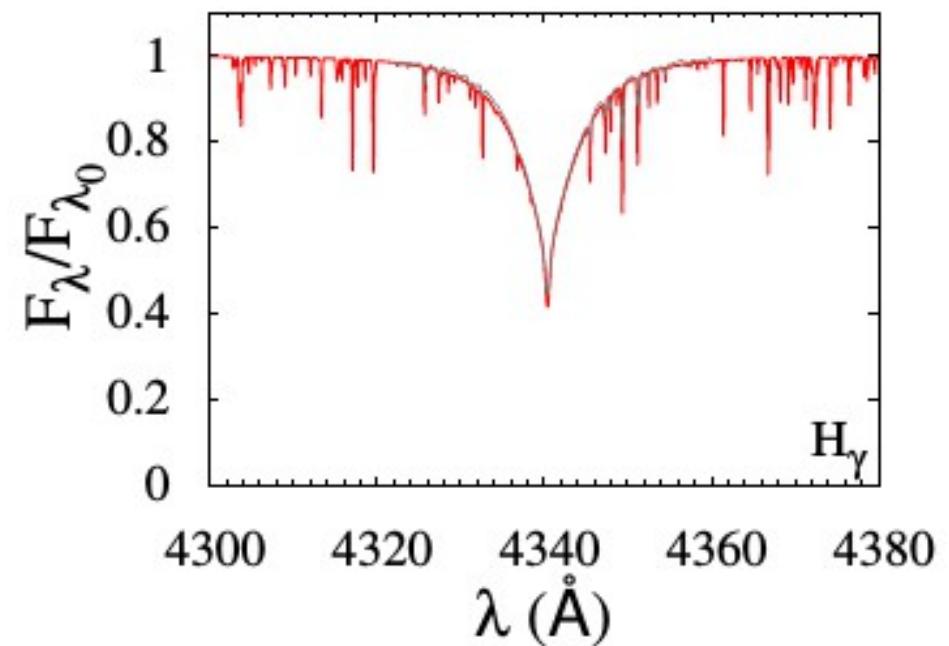
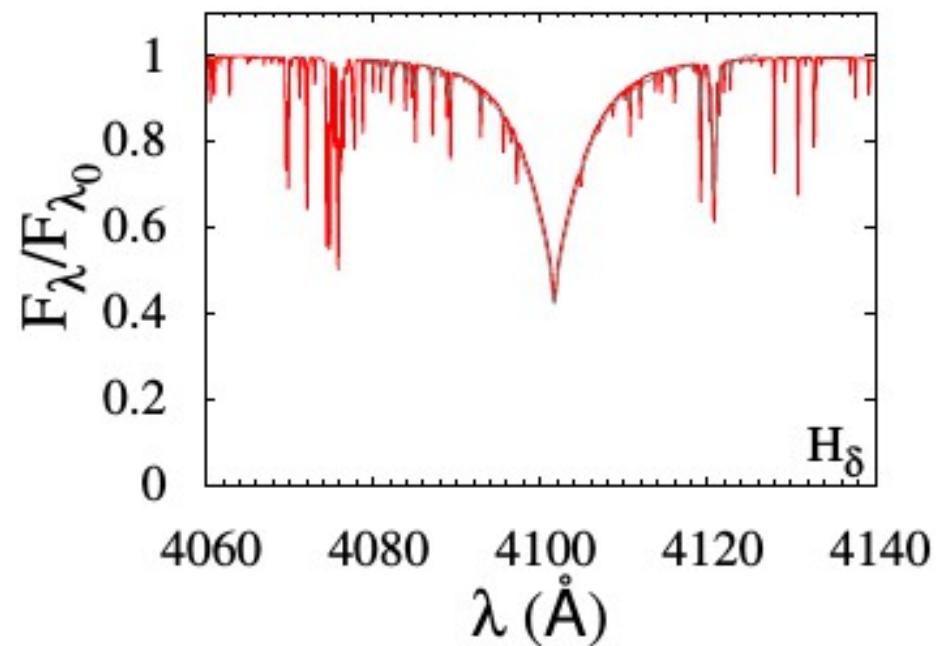
Teff,logg belirleme

- Fotometrik Kalibrasyon+
- Balmer Seri çizgileri+
- İyonizasyon Dengesi-
- Eksitaston potansiyeli-
- SED (Spectral Energy Dist.)-

Method

- **IRAF, splot,continuum task**
 - Normalizasyon
 - Gauss,Fit, EW
- Atomik veriler, **VALD**
- **SYNPLOT, Tlusty**
 - Teff
 - Logg
 - Vsini
 - Microtürbülans Hızı
 - CNO bolluk





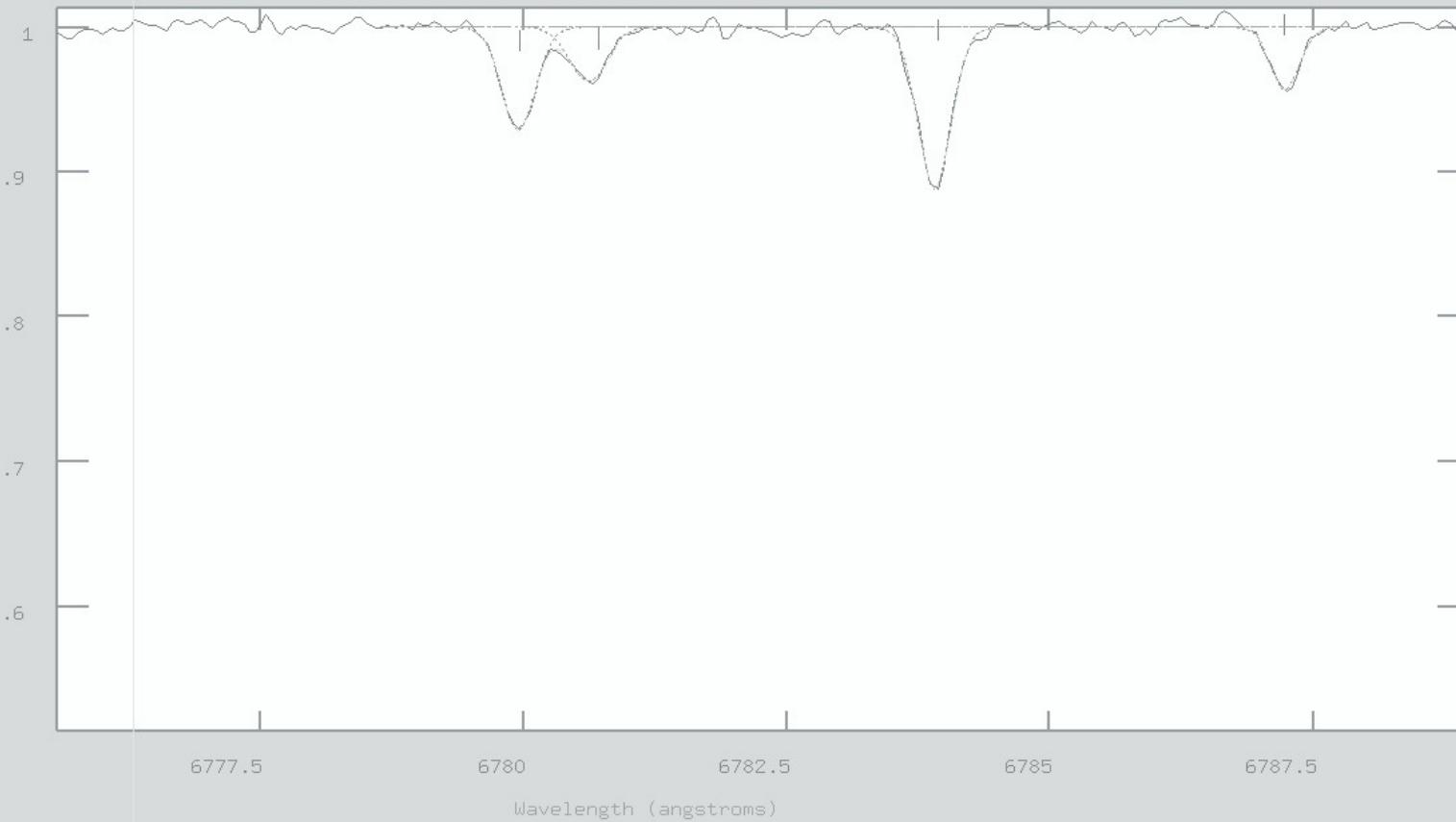
T_{eff} (K)	$\log g$ (cgs)	Kaynak
22000	4.00	Nieva (2013)
20700	3.81	Koleva & Vazdekis (2011)
20454	3.79	Prugniel P. (2011)
22670	4.02	Gies & Lambert (1992)
21913	3.70	Peters (1976)
20160	...	Snijders (1969)

λ (Å)	T_{eff} (K)	$\log g$ (cgs)	χ^2
4100	22050	4.10	0.0322
4300	22650	3.90	0.0414
4800	23350	3.90	0.0241
ort.	~ 22700	~ 3.97	
sd.	~ 900	~ 0.10	

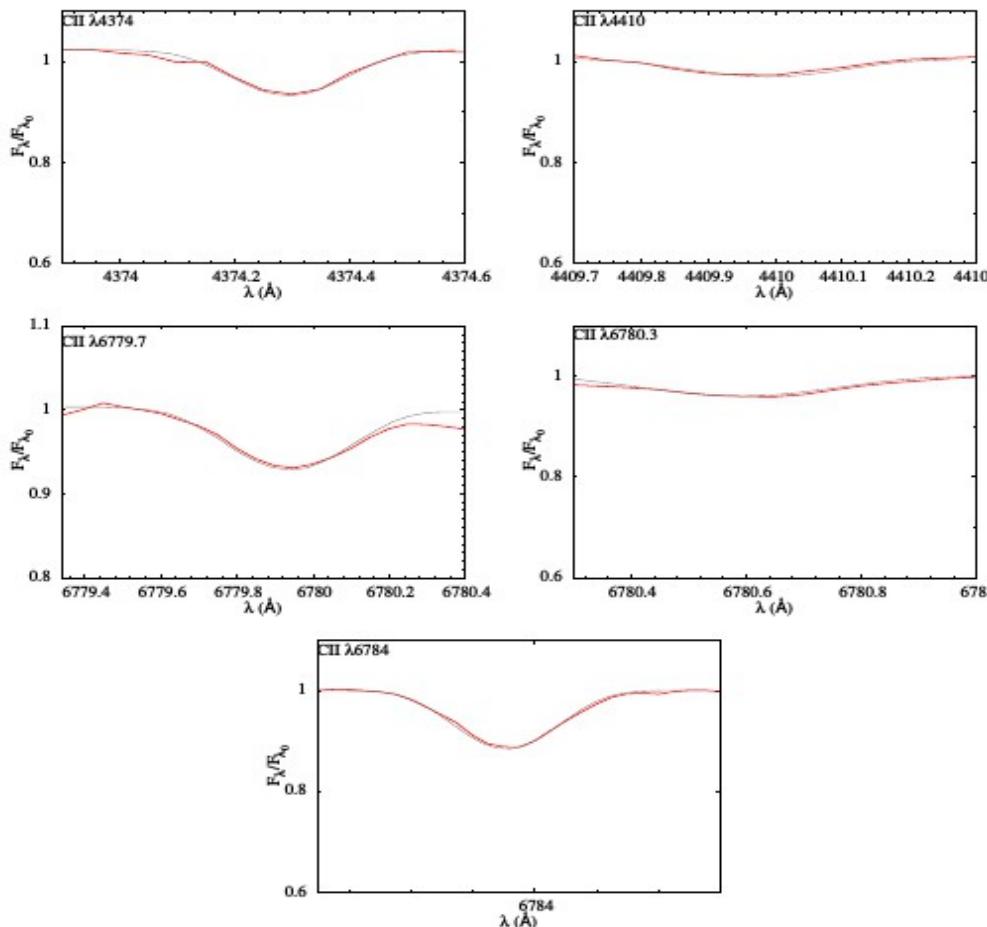
	T_{eff} (K)	$\log g$ (cgs)
Q	20454	3.79
[u-b]	20700	3.81
[c1]	22670	4.02
ort.	23000	4.00
sd.	1000	0.10

λ (Å)	He/H
4009	0.12
4026	0.12
4388	0.15
4471	0.11
4713	0.09
Average	0.12
Std.Dev.	0.02

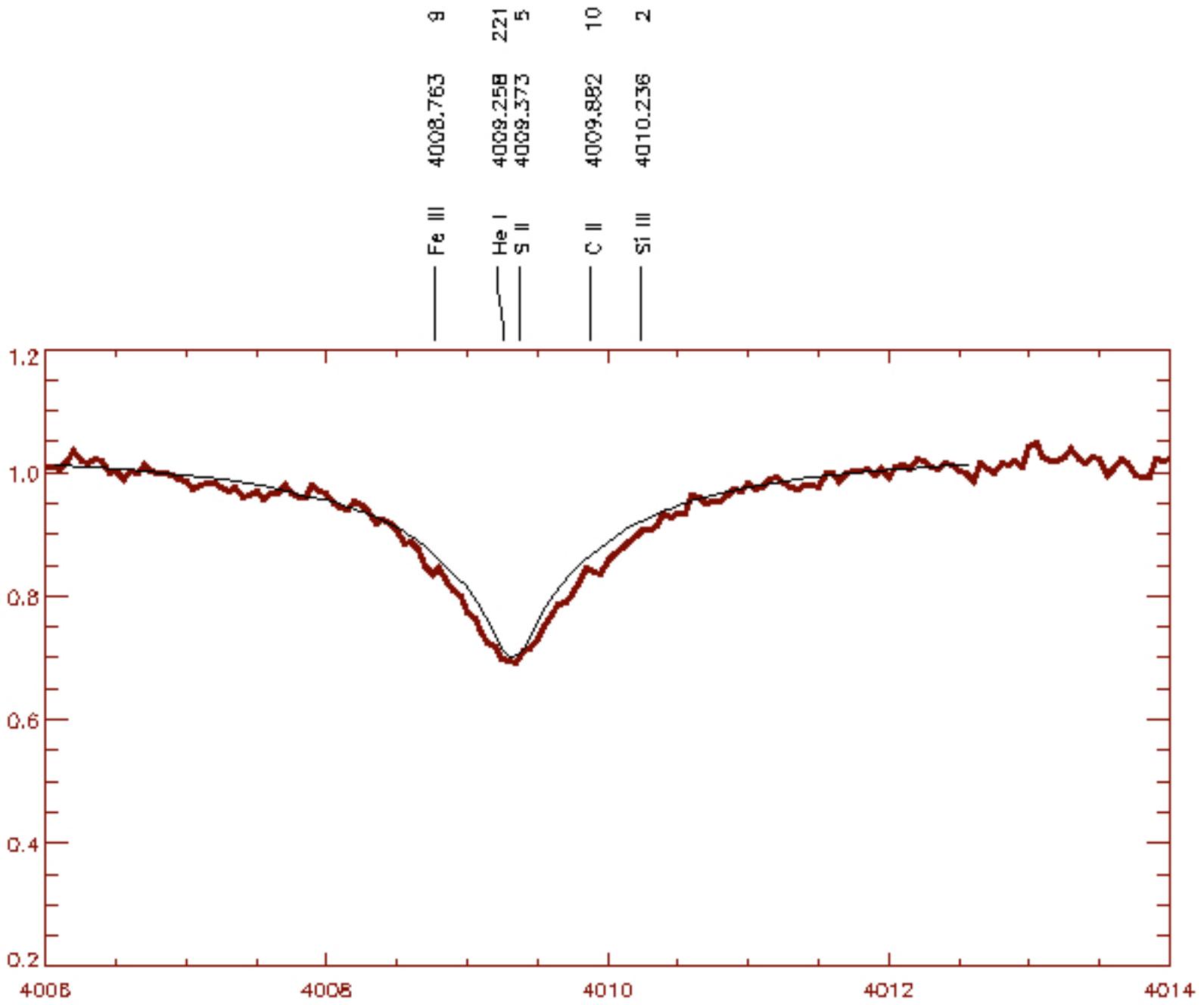
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[535nm.fits]: HD000886 120. ap:1 beam:1

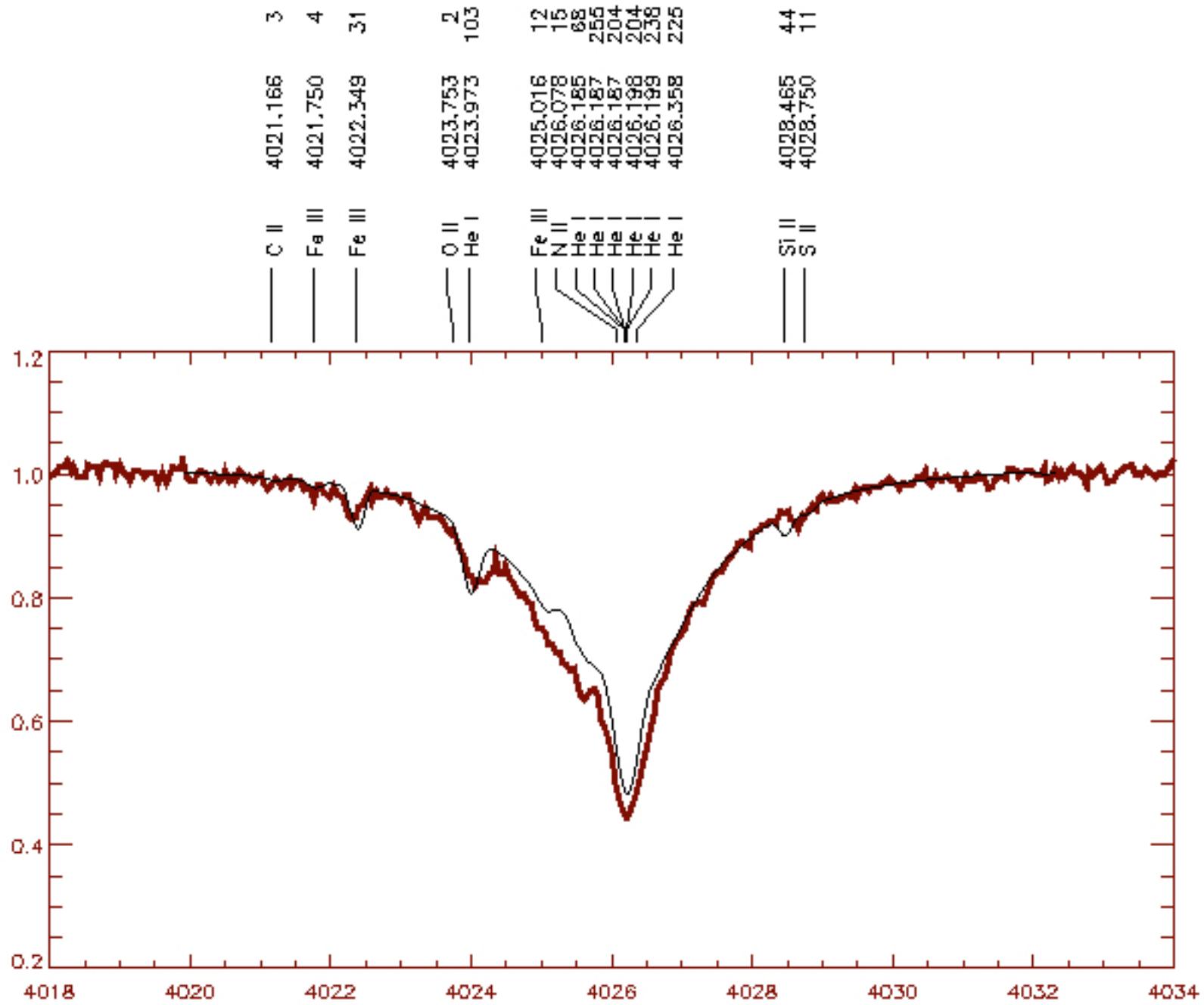


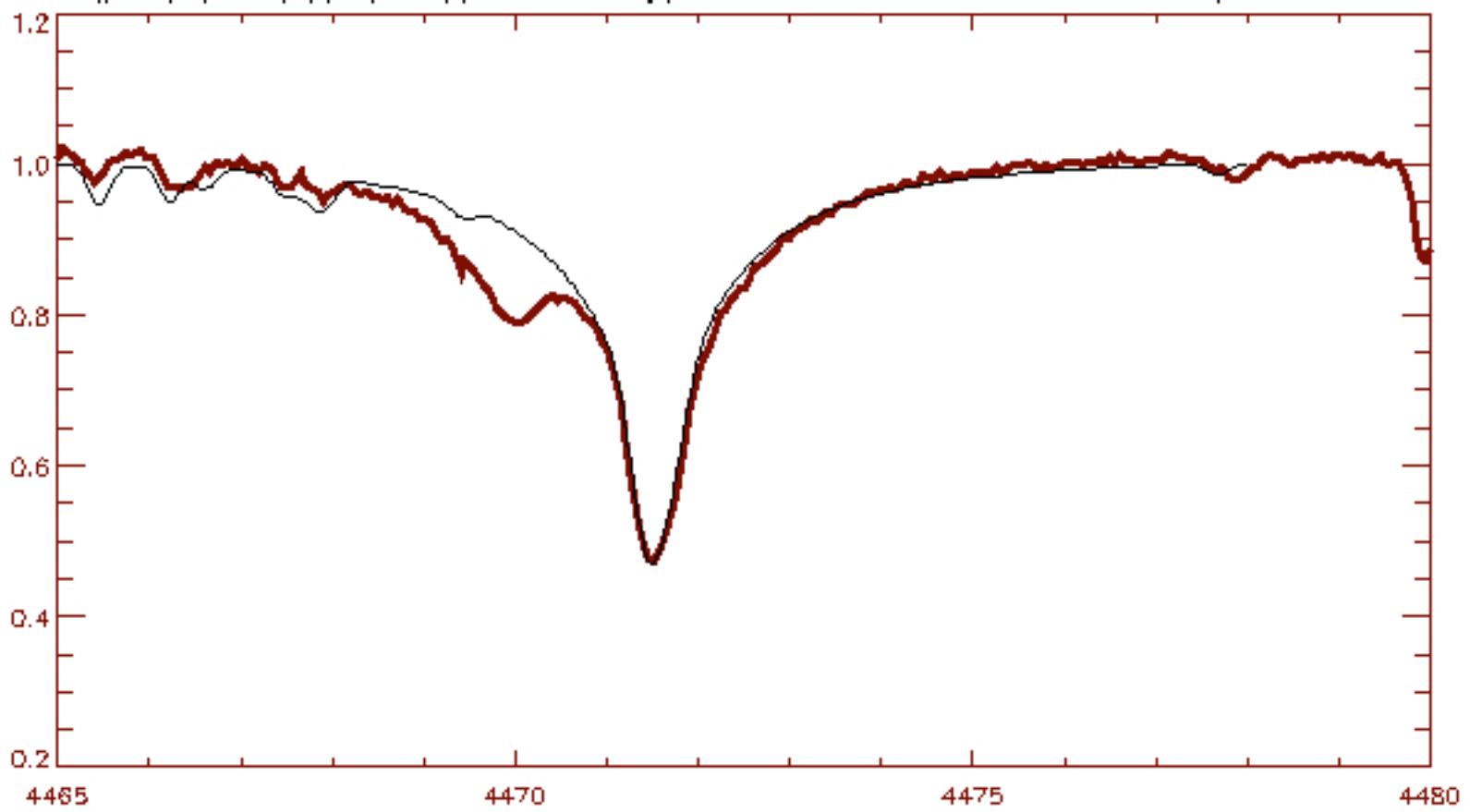
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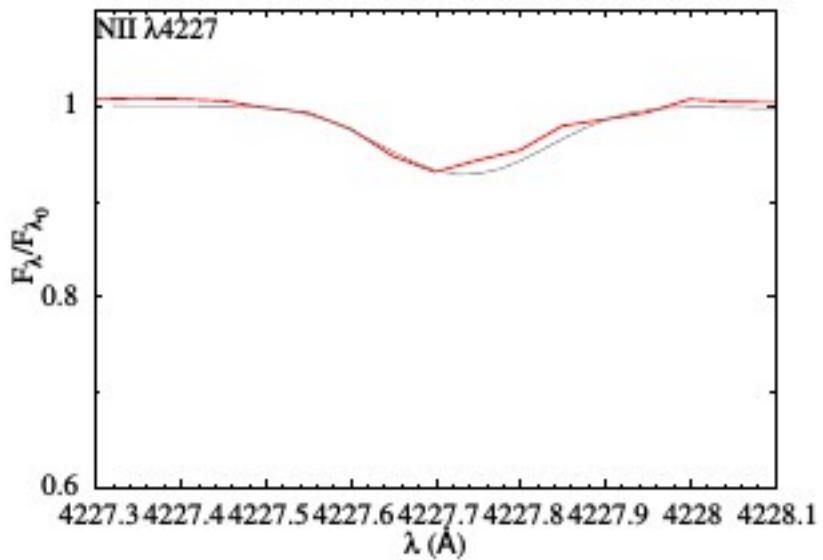
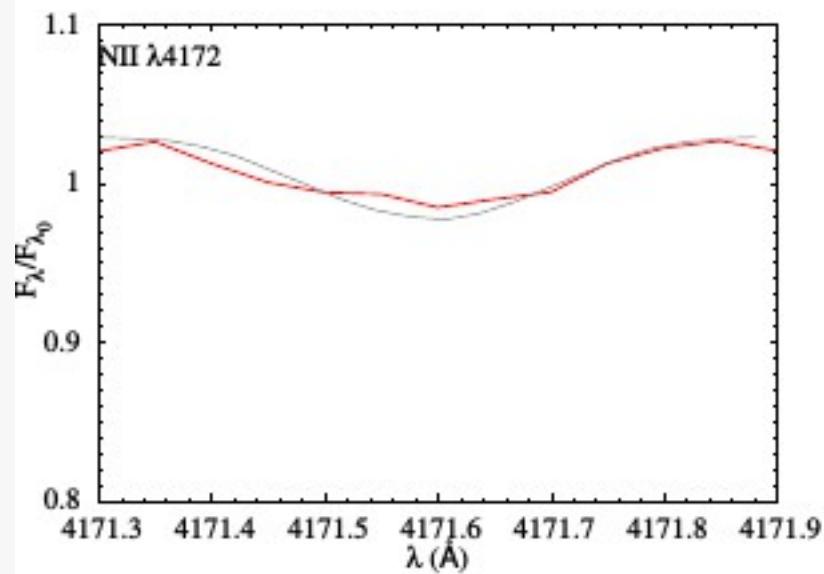
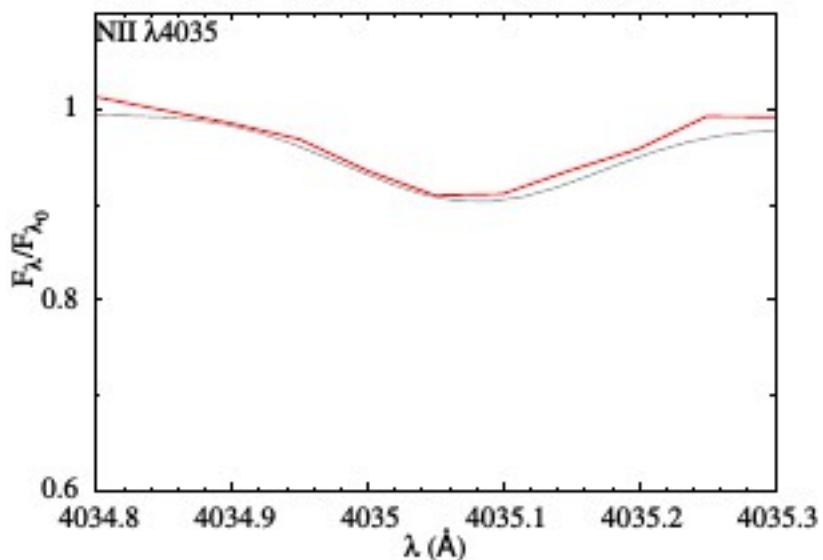
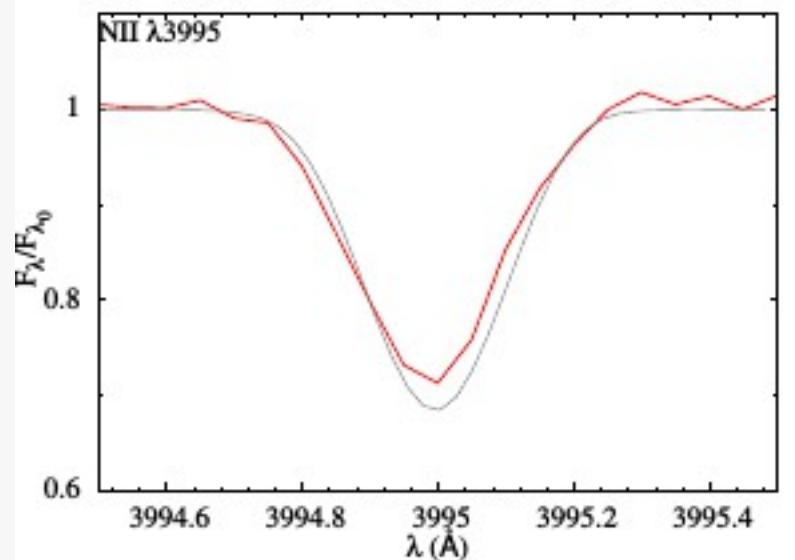


Species	Multiplet	λ (mÅ)	log gf	EP(ev)	W_λ (mÅ)	NLTE	LTE	$\Delta \log \epsilon$
						log N/N _T	log N/N _T	
C II(8.52)						8.23±0.25	8.23±0.23	
	42	4374.27	0.66	24.50	21.4	8.23	8.25	-0.02
	40	4410.06	0.50	24.50	9.7	7.95	7.96	-0.01
	39	4411.20	0.53	24.50	19.5	8.01	8.04	-0.03
		4411.52	0.69	24.50	17.2	8.00	8.02	-0.02
	14	6779.74	0.03	20.61	29.6	8.40	8.39	0.01
		6780.27	-0.38	20.61	17.4	8.53	8.48	+0.05
		6783.75	0.30	20.62	46.7	8.41	8.48	-0.07

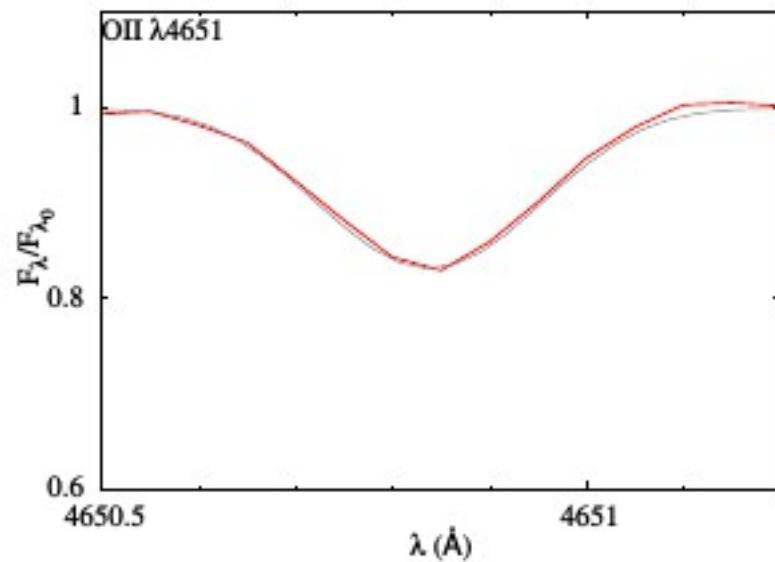
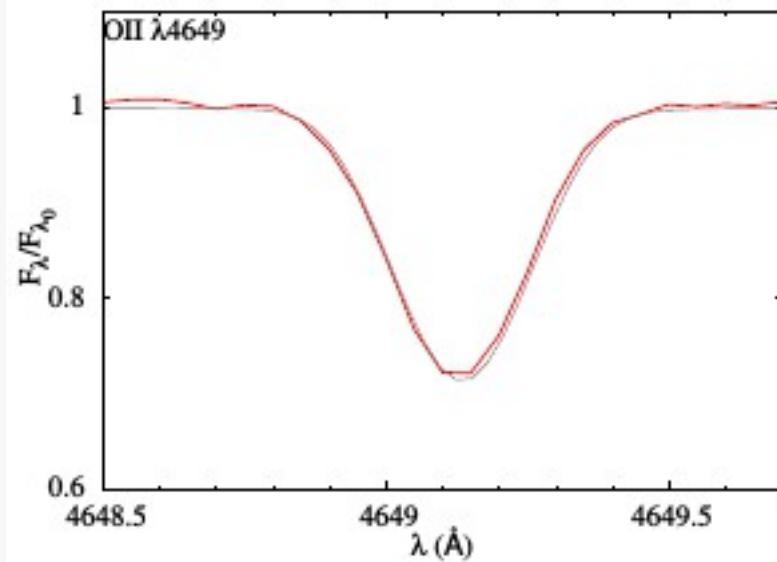
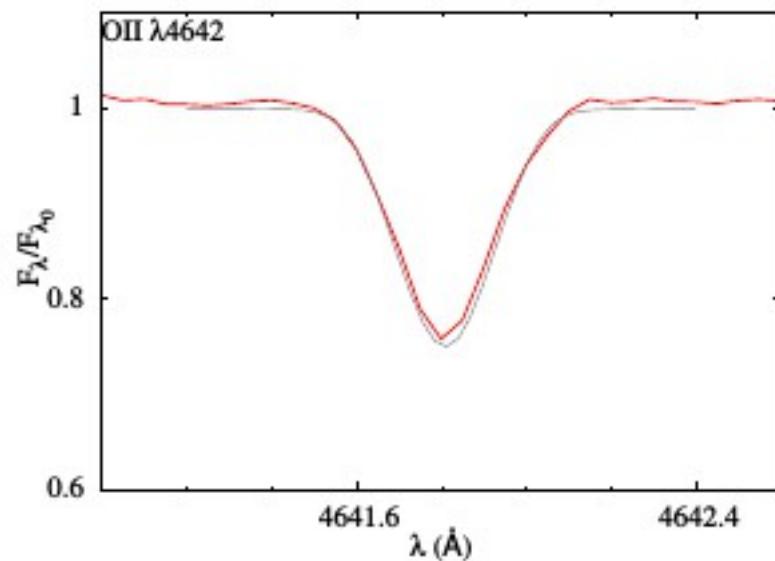
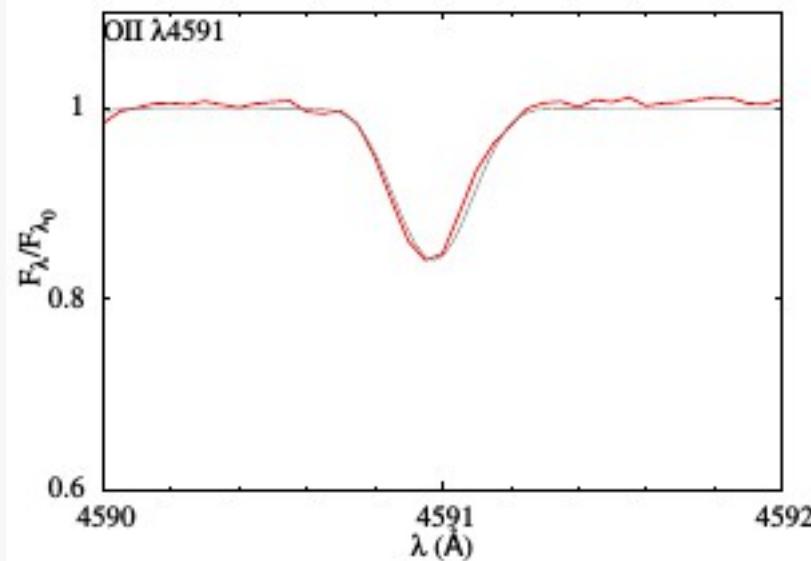








Species	Multiplet	$\lambda(\text{m}\text{\AA})$	log gf	EP(ev)	$W_\lambda(\text{m}\text{\AA})$	NLTE	LTE	$\Delta \log \epsilon$
						log N/N _T	log N/N _T	
N II(7.92)						7.75±0.14	7.72±0.15	
	12	3995.00	0.21	18.42	79.0	7.78	7.86	-0.01
	39	4035.08	0.69	23.02	26.1	7.53	7.47	0.06
	43	4171.63	0.28	23.10	13.3	7.72	7.63	0.09
	42	4176.16	0.60	23.10	17.9	7.76	7.70	0.06
	33	4227.75	-0.06	21.51	17.0	7.78	7.72	0.06
	48	4236.93	0.38	23.14	32.8	7.66	7.63	0.03
	21	4507.56	-0.82	20.58	5.5	7.77	7.74	0.03
	20	4779.71	-0.59	20.56	10.2	7.85	7.74	0.11
	20	4788.13	-0.36	20.56	14.6	7.82	7.82	0.00



Species	Multiplet	$\lambda(\text{m}\text{\AA})$	$\log \text{gf}$	EP(ev)	$W_\lambda(\text{m}\text{\AA})$	NLTE	LTE	$\Delta \log \epsilon$
						$\log N/N_T$	$\log N/N_T$	
O II(8.83)						8.67±0.08	8.75±0.08	
	6	3982.71	-0.70	23.34	28.2	8.67	8.71	-0.04
	10	4078.84	-0.09	25.53	25.5	8.73	8.75	-0.02
	2	4345.56	-0.70	22.88	25.3	8.76	8.83	0.00
	5	4452.38	-0.79	23.34	29.0	8.73	8.79	0.00
	15	4590.97	-0.09	25.55	44.6	8.59	8.69	-0.10
	49	4071.24	-0.28	28.57	13.8	8.83	8.83	0.00
	36	4185.56	0.60	28.24	27.3	8.57	8.83	0.00
	1	4641.84	0.06	22.88	71.1	8.67	8.84	-0.17
		4649.14	0.31	22.90	86.5	8.60	8.84	-0.20
		4650.84	-0.36	22.87	47.3	8.60	8.65	-0.15
		4661.63	-0.28	22.88	52.4	8.67	8.69	-0.02
		4676.24	-0.39	22.90	43.2	8.62	8.73	-0.11
	25	4696.35	-1.38	22.90	13.0	8.73	8.77	-0.04
		4705.35	-0.28	26.14	40.3	8.62	8.69	-0.07
	24	4703.16	0.26	28.39	14.3	8.55	8.58	-0.03
		4710.10	-0.23	26.11	20.9	8.70	8.76	-0.00
		4751.24	-0.28	26.14	7.8	8.69	8.72	-0.03
	28	4890.86	-0.44	26.19	14.2	8.79	8.85	-0.06
		4906.83	-0.16	26.19	18.4	8.63	8.71	-0.08
	33	4943.01	0.24	26.45	24.9	8.55	8.62	-0.00
		4955.71	0.57	26.45	9.5	8.74	8.78	-0.04

Table 6. Metal abundances for the sample stars.

#	HD	C	N	O	Ne	Mg	Si	Fe
1	36591	8.33±0.08 (30)	7.75±0.09 (61)	8.75±0.11 (53)	8.09±0.08 (21)	7.58±0.10 (6)	7.50±0.04 (5)	7.48±0.11 (32)
2	61068	8.27±0.07 (23)	8.00±0.12 (61) ^a	8.76±0.09 (49)	8.07±0.11 (17)	7.56±0.03 (3)	7.53±0.06 (5)	7.51±0.11 (28)
3	63922	8.33±0.07 (19)	7.77±0.08 (23)	8.79±0.10 (39)	8.07±0.07 (8)	7.60±0.01 (2)	7.49±0.12 ^b	7.51±0.08 (9)
4	74575	8.37±0.10 (19)	7.92±0.10 (56) ^a	8.79±0.08 (45)	8.05±0.08 (12)	7.51±0.10 (6)	7.52±0.12 ^b	7.51±0.09 (27)
5	122980	8.32±0.09 (22)	7.76±0.08 (47)	8.72±0.06 (52)	8.07±0.07 (14)	7.50±0.05 (4)	7.25±0.04 (4) ^a	7.44±0.11 (27)
6	149438	8.30±0.12 (32)	8.16±0.12 (73) ^a	8.77±0.08 (49)	8.14±0.07 (18)	7.62±0.03 (3)	7.52±0.06 (2)	7.54±0.09 (21)
7	886	8.37±0.08 (17)	7.76±0.07 (40)	8.73±0.11 (52)	8.11±0.08 (13)	7.61±0.05 (4)	7.38±0.03 (5)	7.51±0.07 (30)